

C O I N ' S



Dimensionality Reduction And Clustering
for Unsupervised Learning in Astronomy



DRACULA

Dimensionality Reduction And Clustering for Unsupervised Learning in Astronomy

arXiv:1512.06810v1

Sasdelli et al., MNRAS, 2016, 461, Issue 2, p.2044-2059

[ascl:1512.009]

Emille E. O. Ishida

*Laboratoire de Physique Corpusculaire - Université Clermont-Auvergne
Clermont Ferrand, France*



Type Ia Sne can be used as standard candles



Type Ia Sne can be used as standard candles

The expansion of the Universe
is accelerating!



Type Ia Sne can be used as standard candles

The expansion of the Universe
is accelerating!



2011



Photo: Lawrence Berkeley National Lab

Saul Perlmutter



Photo: Belinda Pratten, Australian National University

Brian P. Schmidt



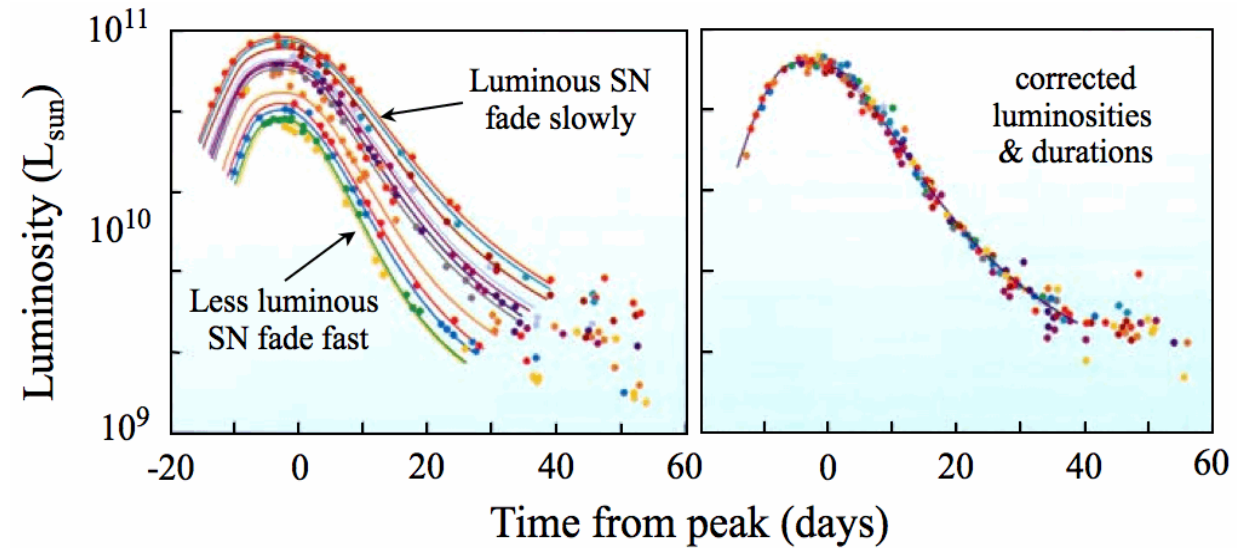
Photo: Scanpix/AFP

Adam G. Riess

The Nobel Prize in Physics 2011 was awarded "for the discovery of the accelerating expansion of the Universe through observations of distant supernovae" with one half to Saul Perlmutter and the other half jointly to Brian P. Schmidt and Adam G. Riess.



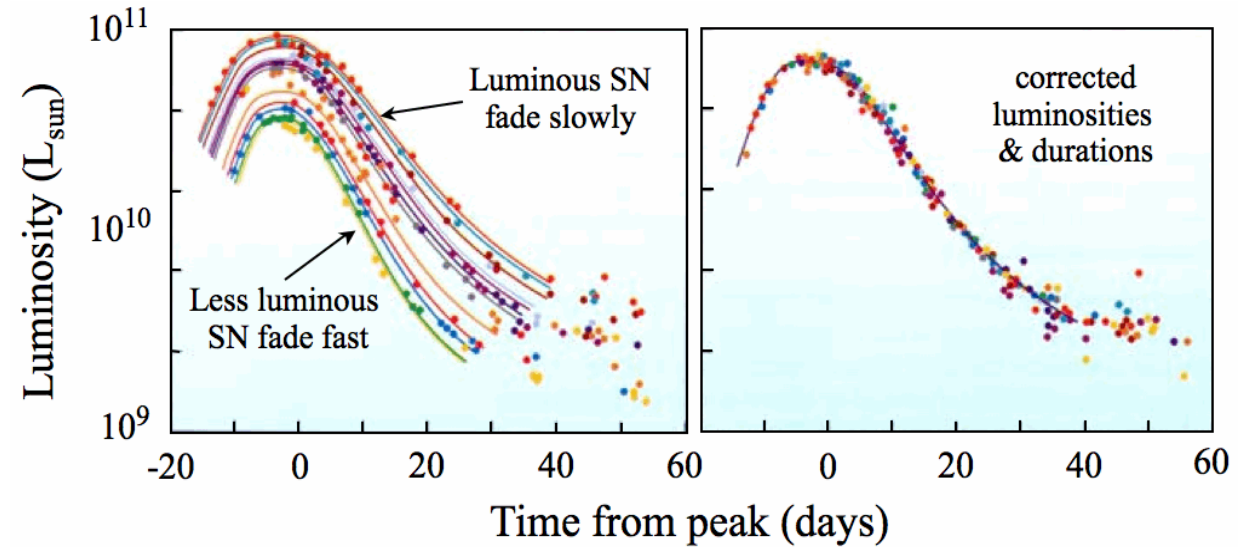
Type Ia SNe are standardizable candles



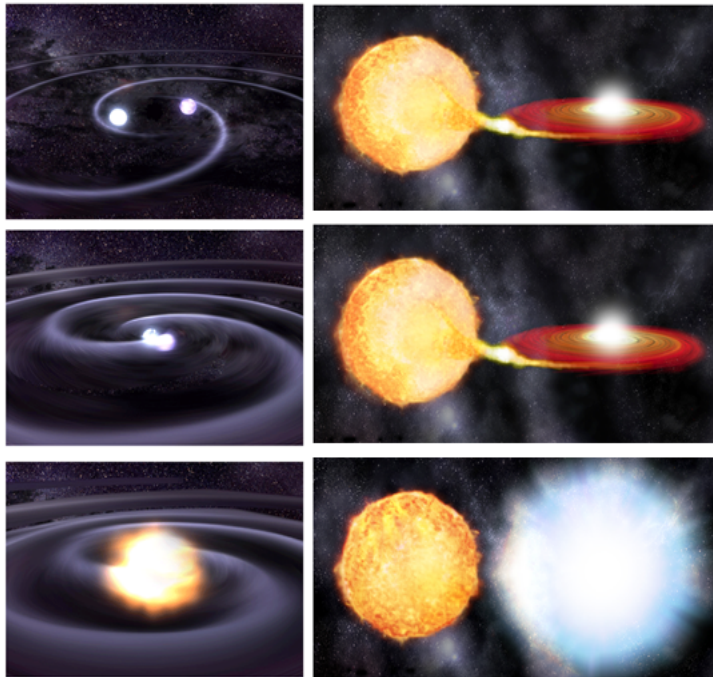
http://community.dur.ac.uk/john.lucey/bridge/SN2015M_bridge.html

Type Ia SNe are standardizable candles

Can we use machine learning to distinguish multiple sub-classes of SN Ia ?

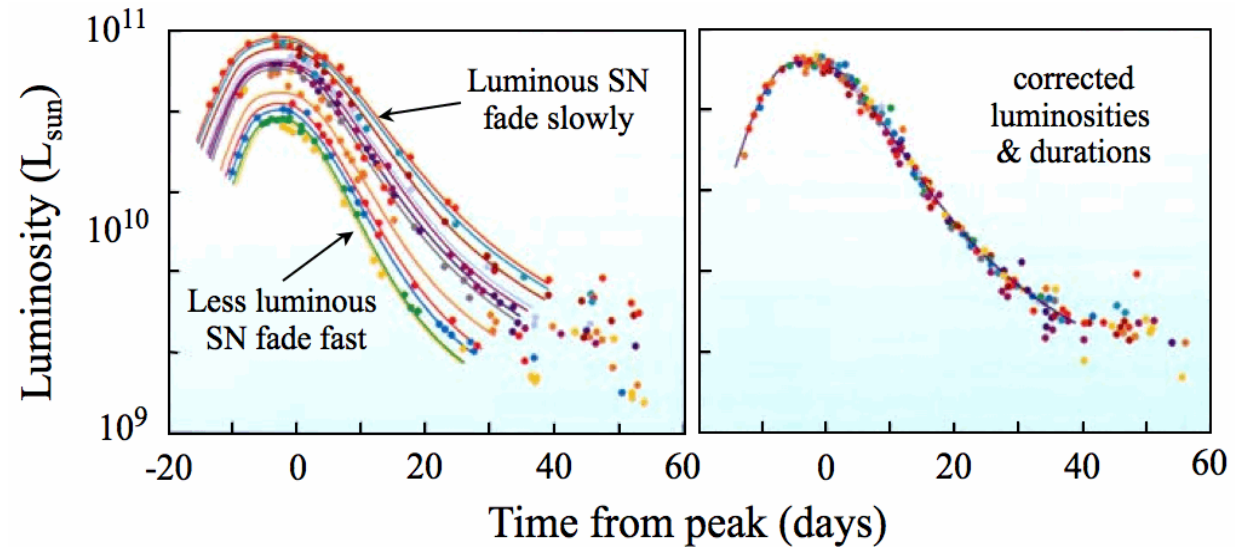


http://community.dur.ac.uk/john.lucey/bridge/SN2015M_bridge.html

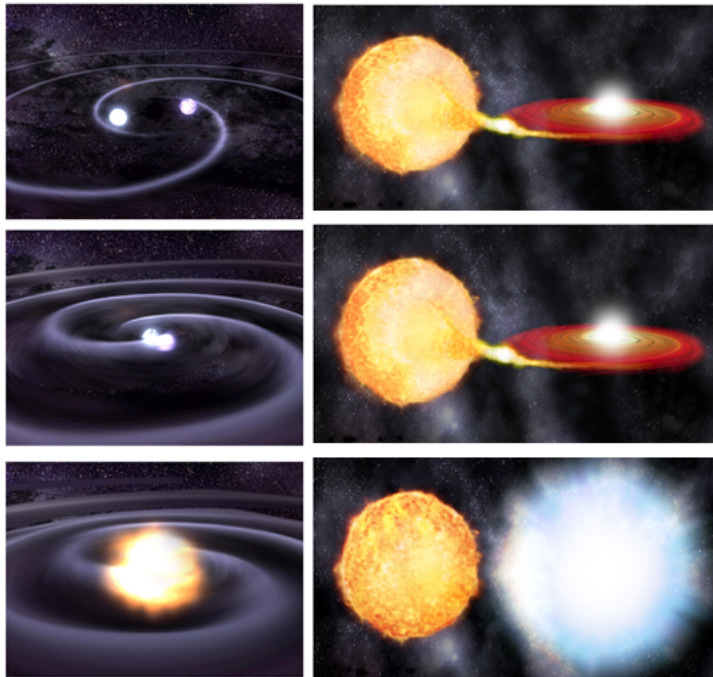


Type Ia SNe are standardizable candles

Can we use machine learning to distinguish multiple sub-classes of SN Ia ?



http://community.dur.ac.uk/john.lucey/bridge/SN2015M_bridge.html



If so, are the resulting classes physically meaningful?

—IDEALLY—

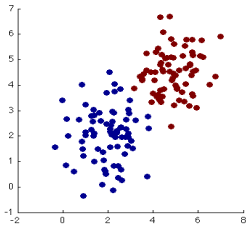


Build data matrix

—IDEALLY—



Build data matrix



Dimensionality Reduction

—IDEALLY—



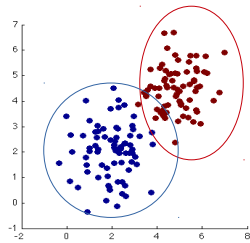
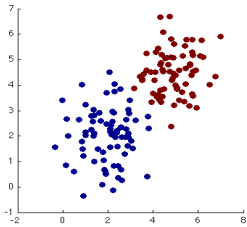
Build data matrix



Dimensionality Reduction



Unsupervised clustering



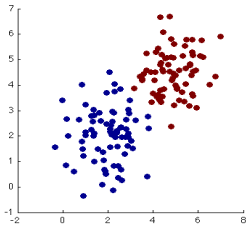
—IDEALLY—



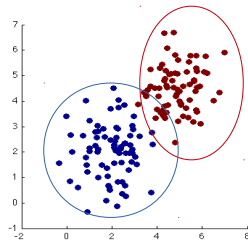
Build data matrix



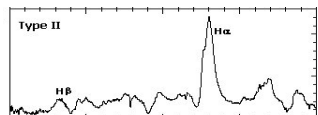
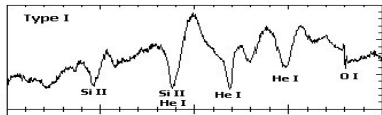
Dimensionality Reduction



Unsupervised clustering



subclasses



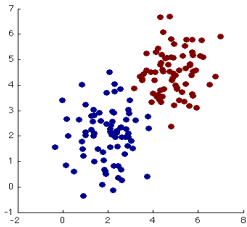
— IDEALLY —



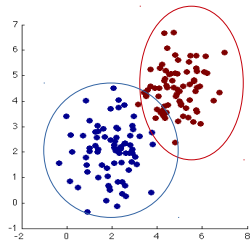
Build data matrix



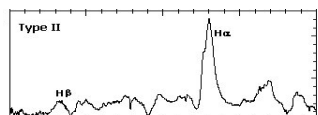
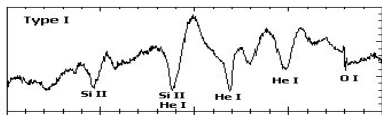
Dimensionality Reduction



Unsupervised clustering



subclasses



~~REALISTICALLY~~

Build data matrix

Problem 1:
Not enough data for training



Dimensionality Reduction



Unsupervised clustering

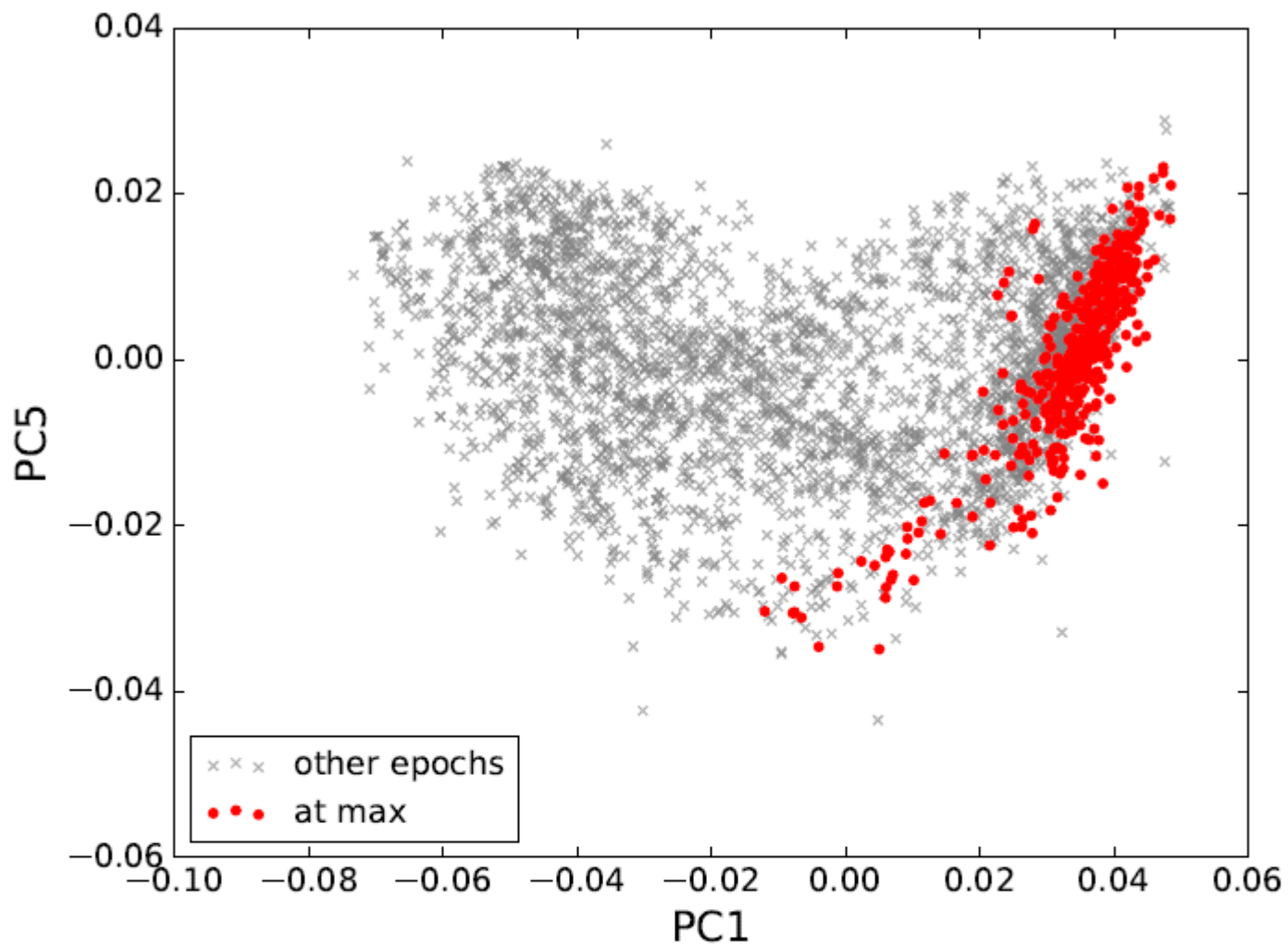


subclasses

PCA parameter space built with all available spectra

even those without an epoch determination

3677 spectra, of which 486 at max



-REALISTICALLY-

Build data matrix

Problem 1:
Not enough data for training
→ Transfer Learning



Dimensionality Reduction

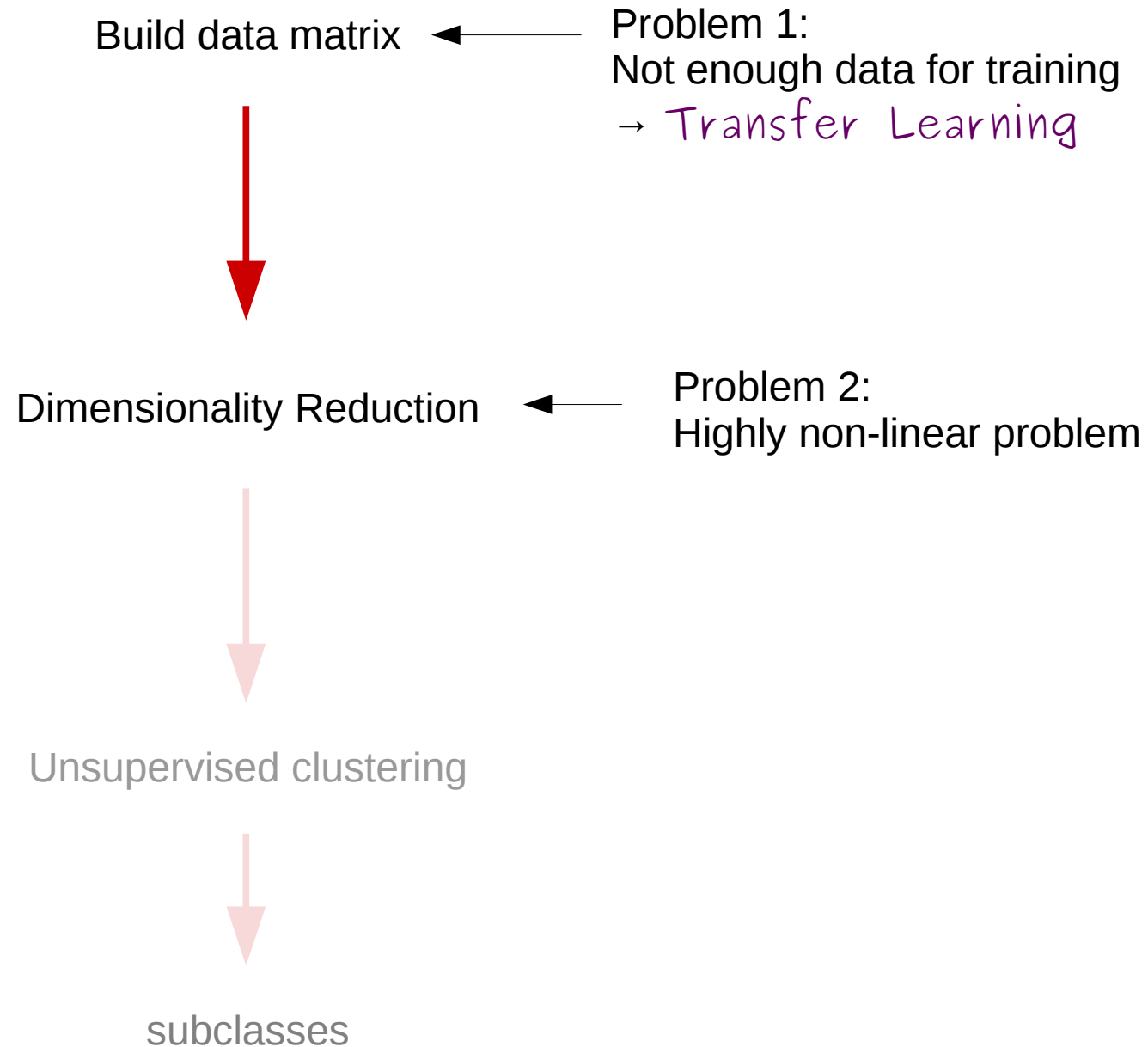


Unsupervised clustering



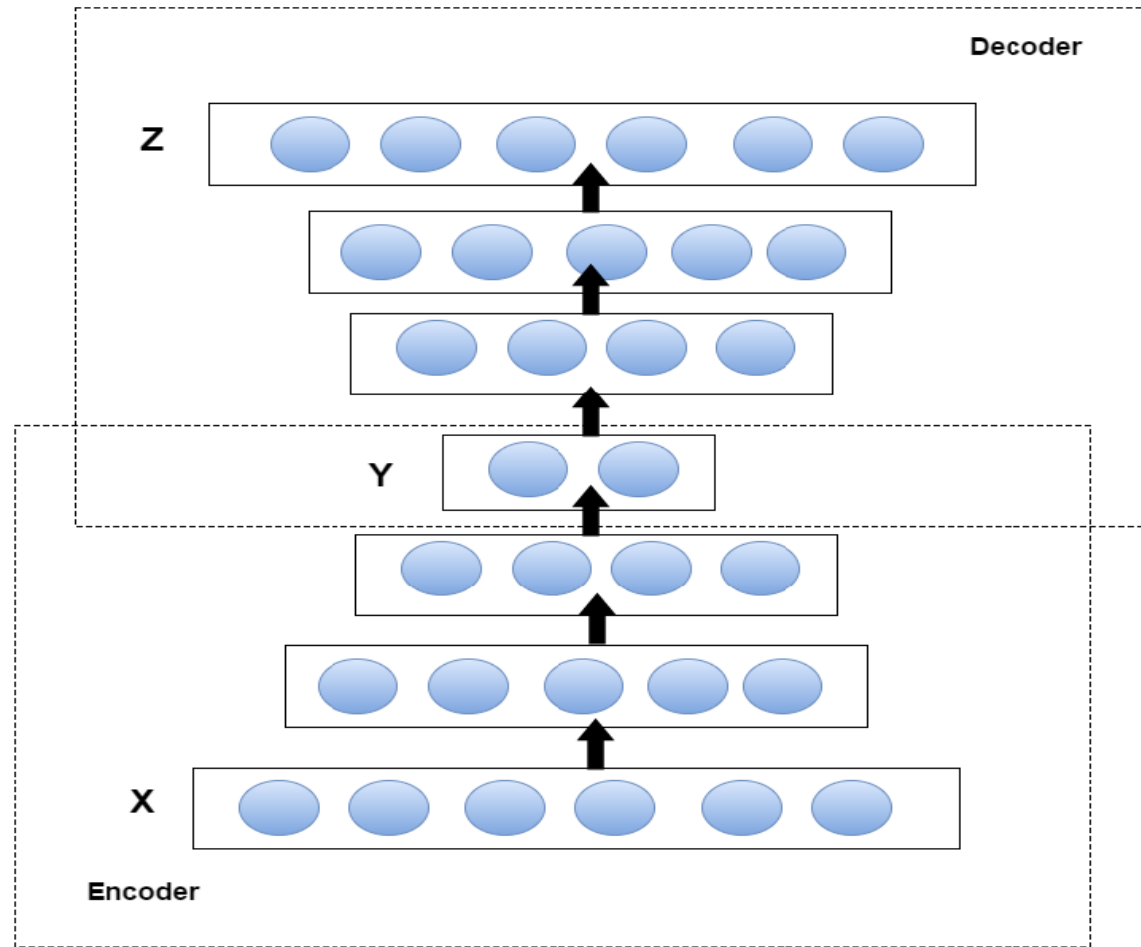
subclasses

-REALISTICALLY-



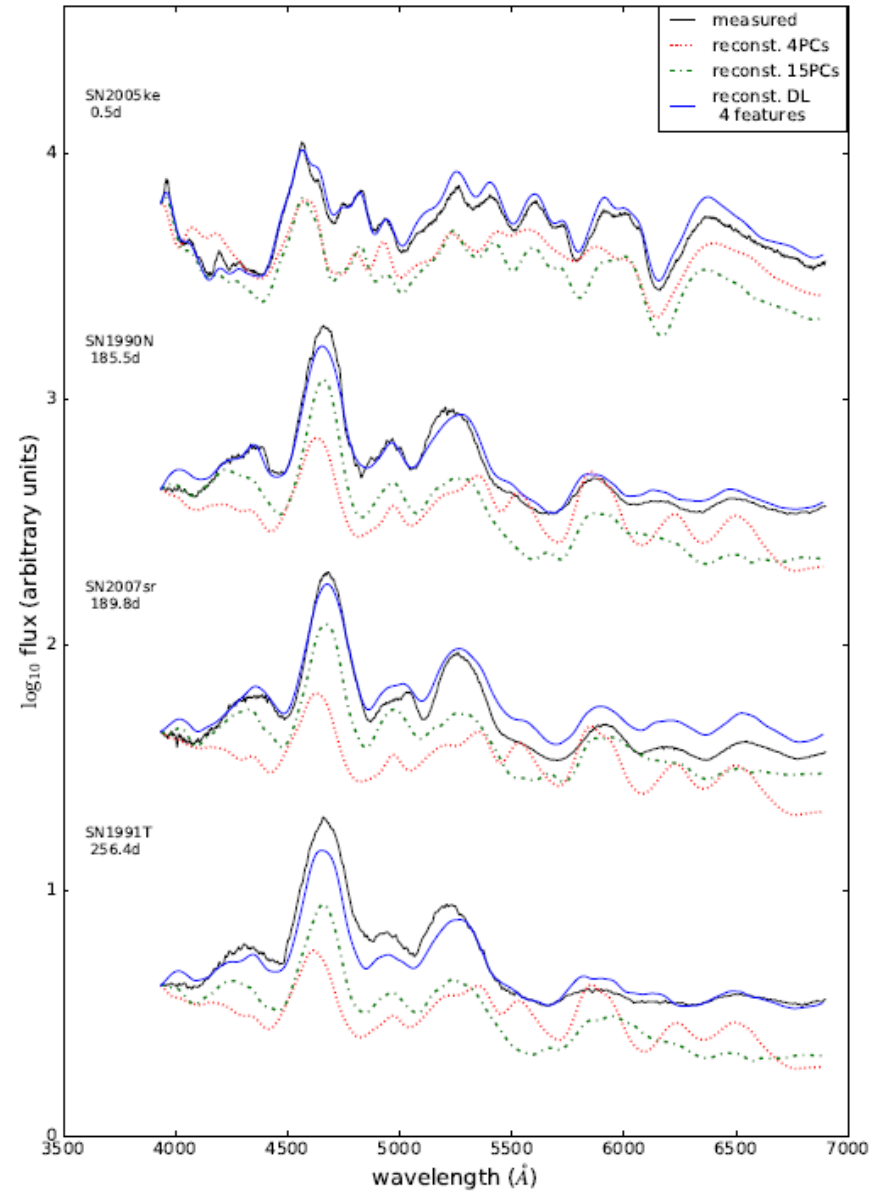
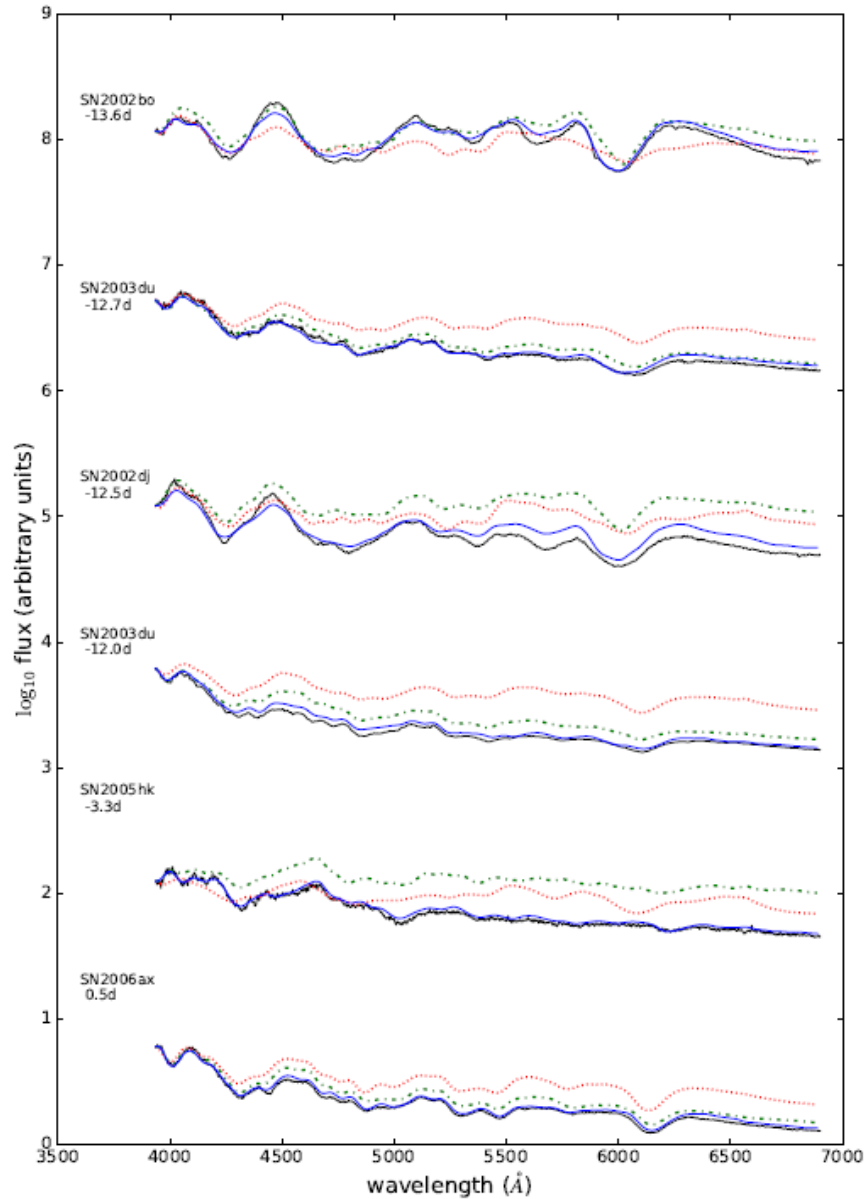
Deep Learning for Dimensionality Reduction

Layers= (120,100,90,50,30,20,4,20,30,50,90,100,120)

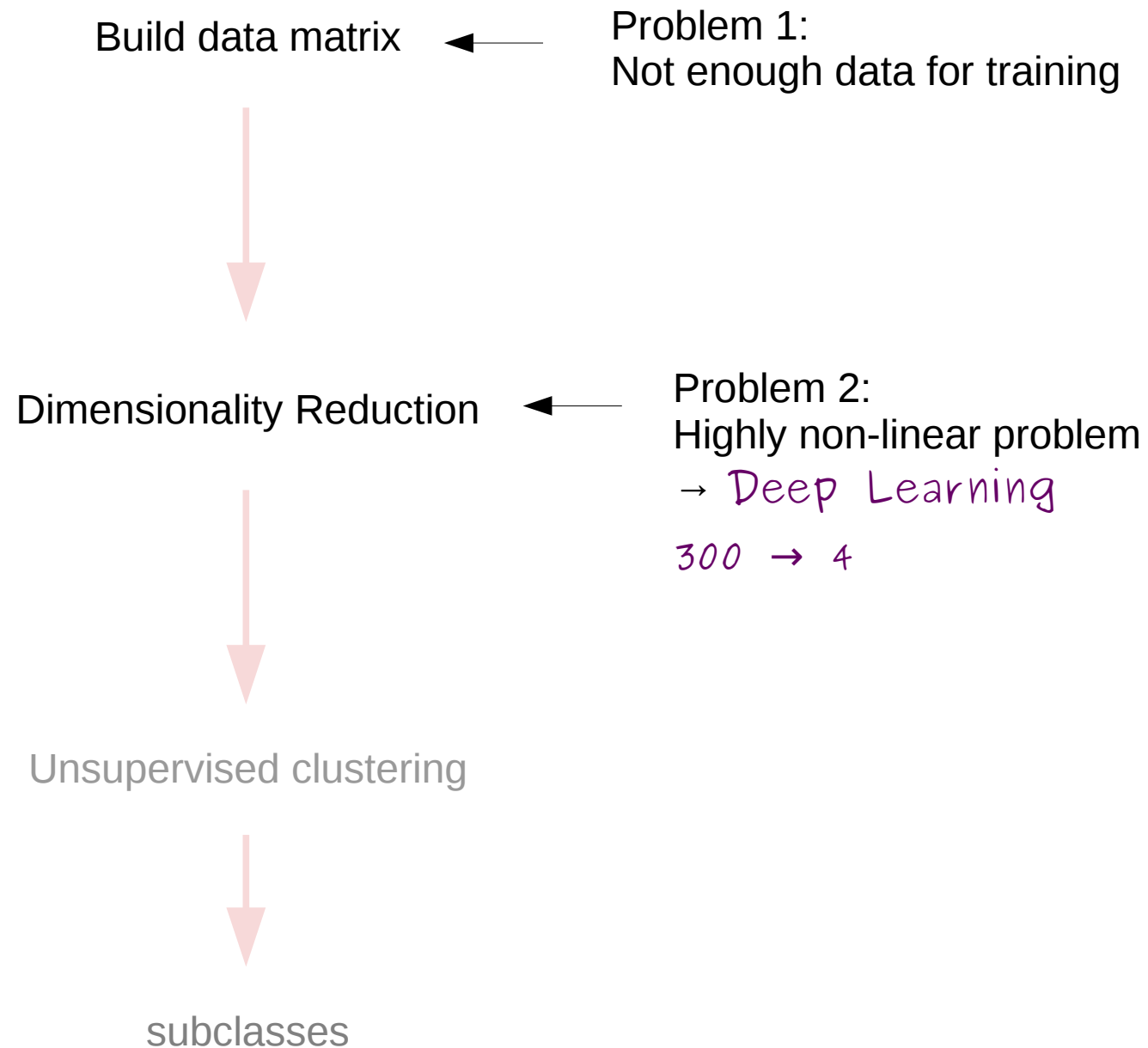


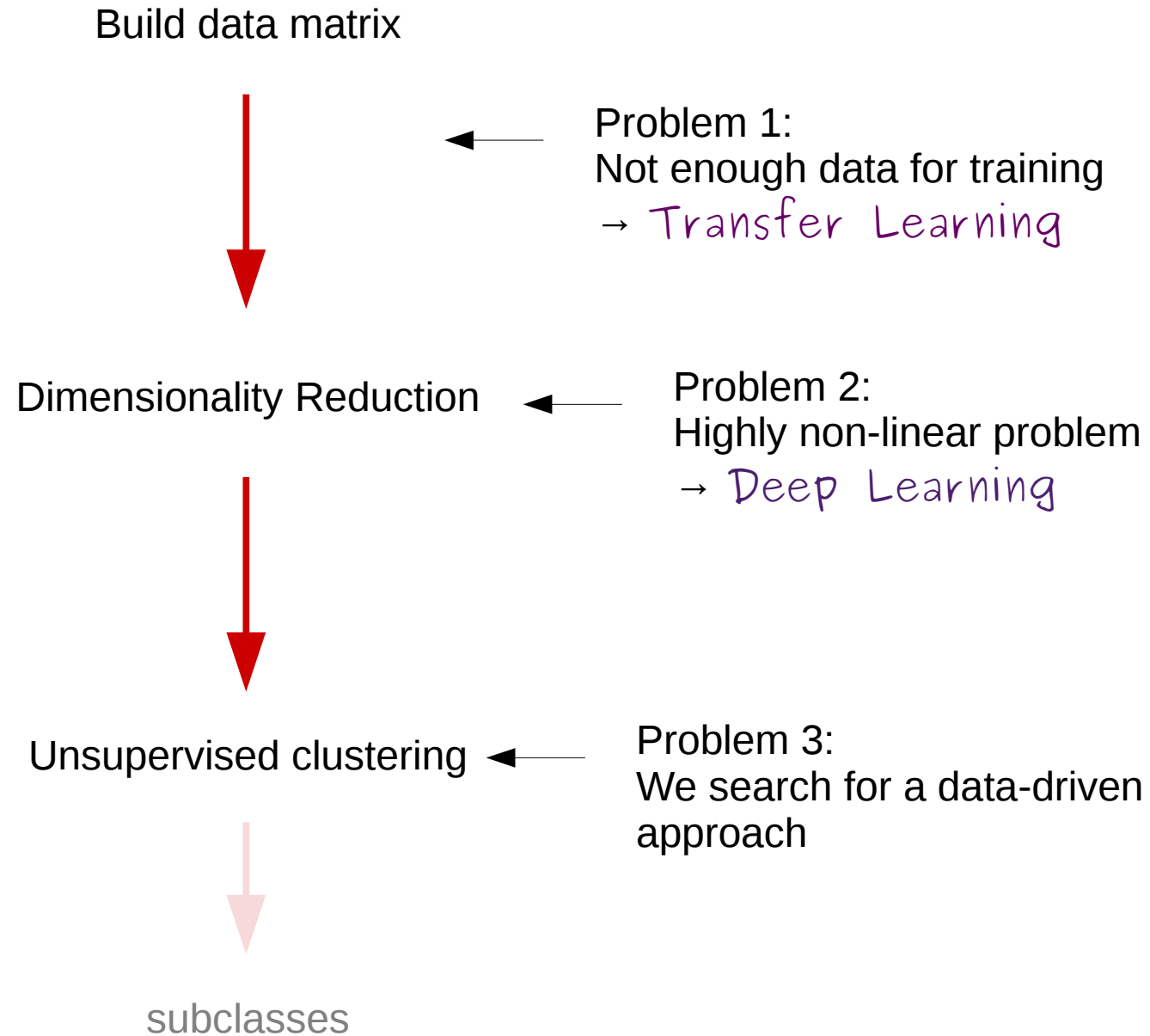
Deep Learning x PCA reconstruction power

DL results are good even at late epochs



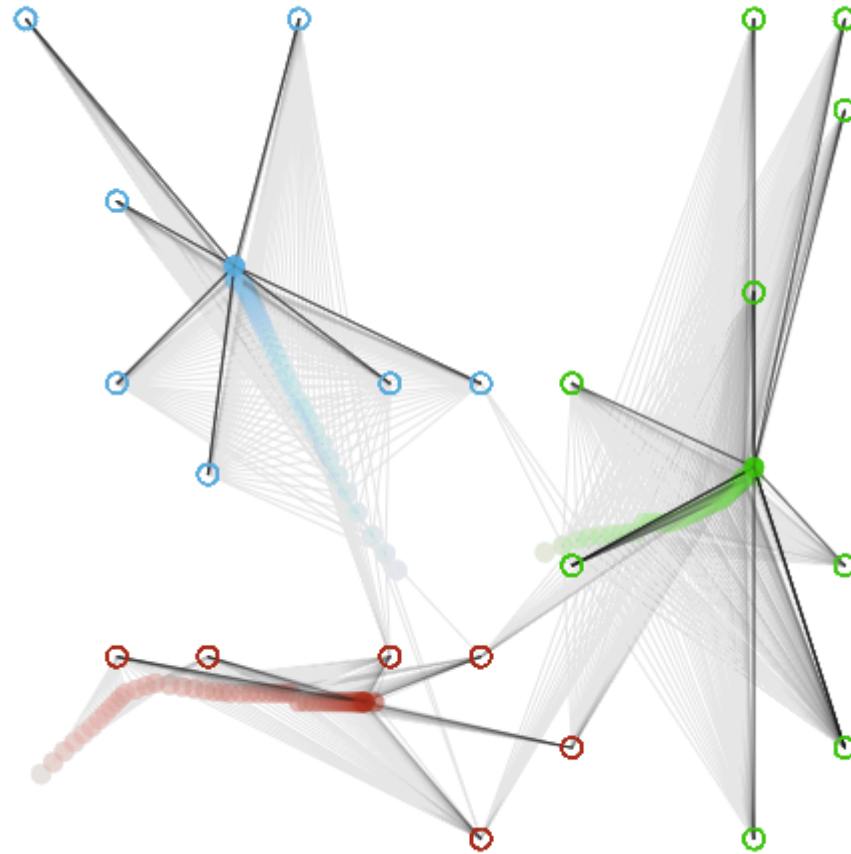
-REALISTICALLY-

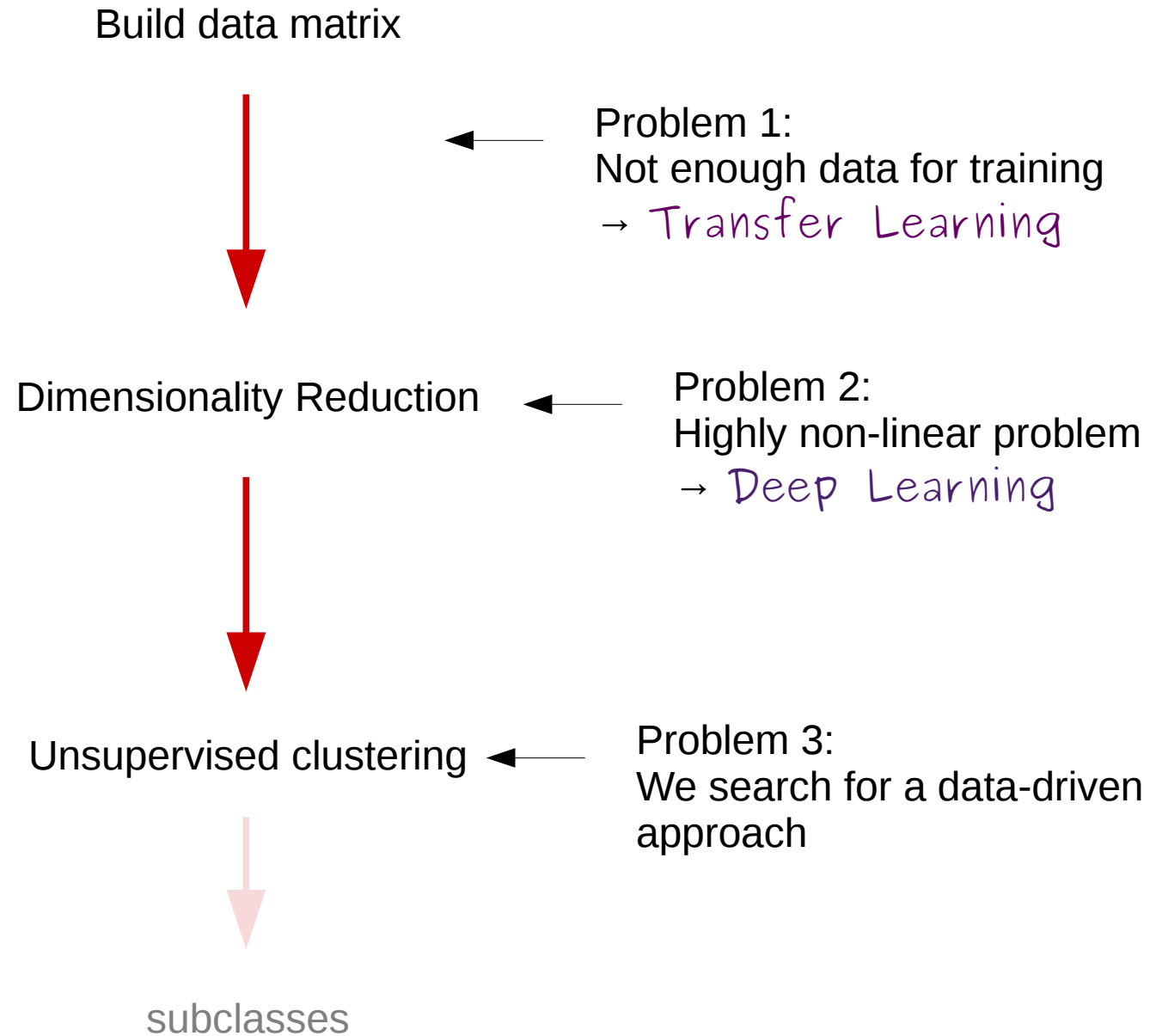




K-means clustering

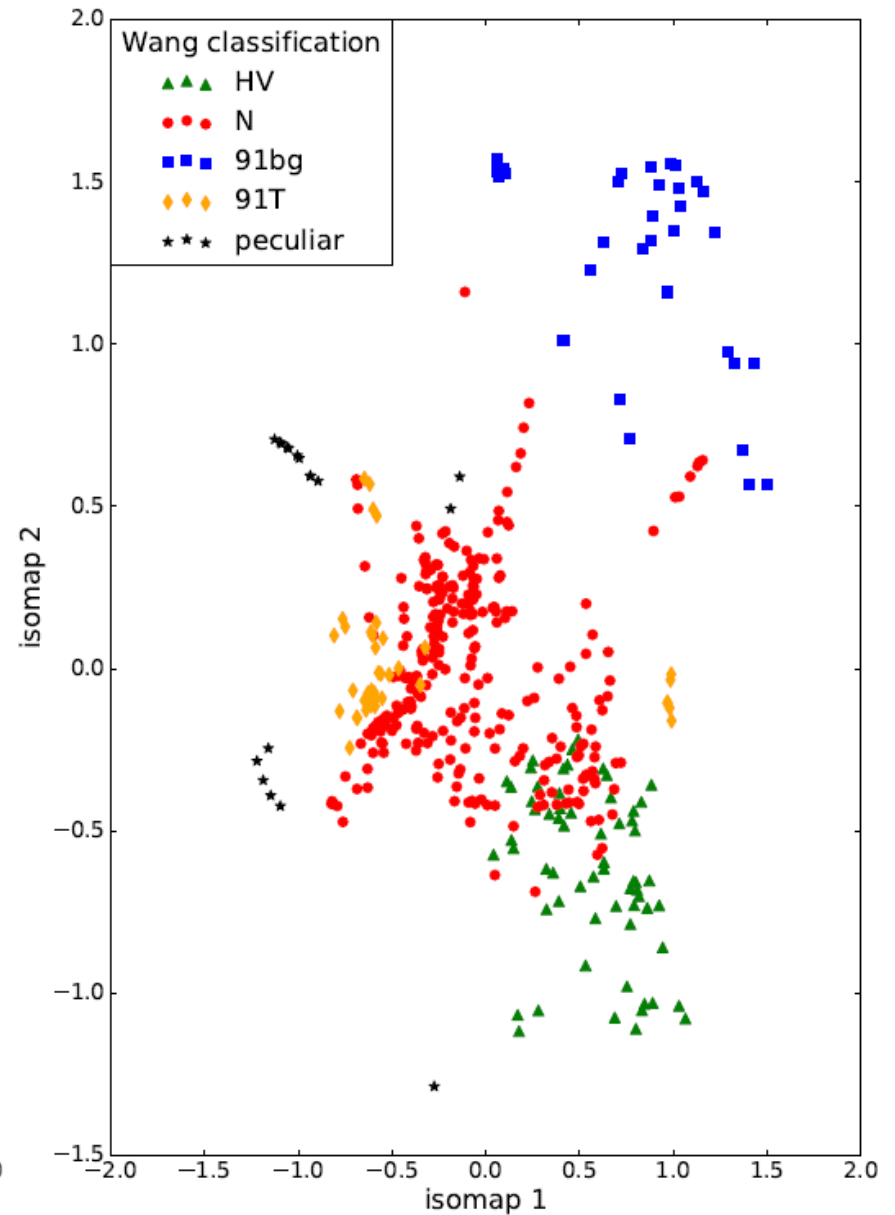
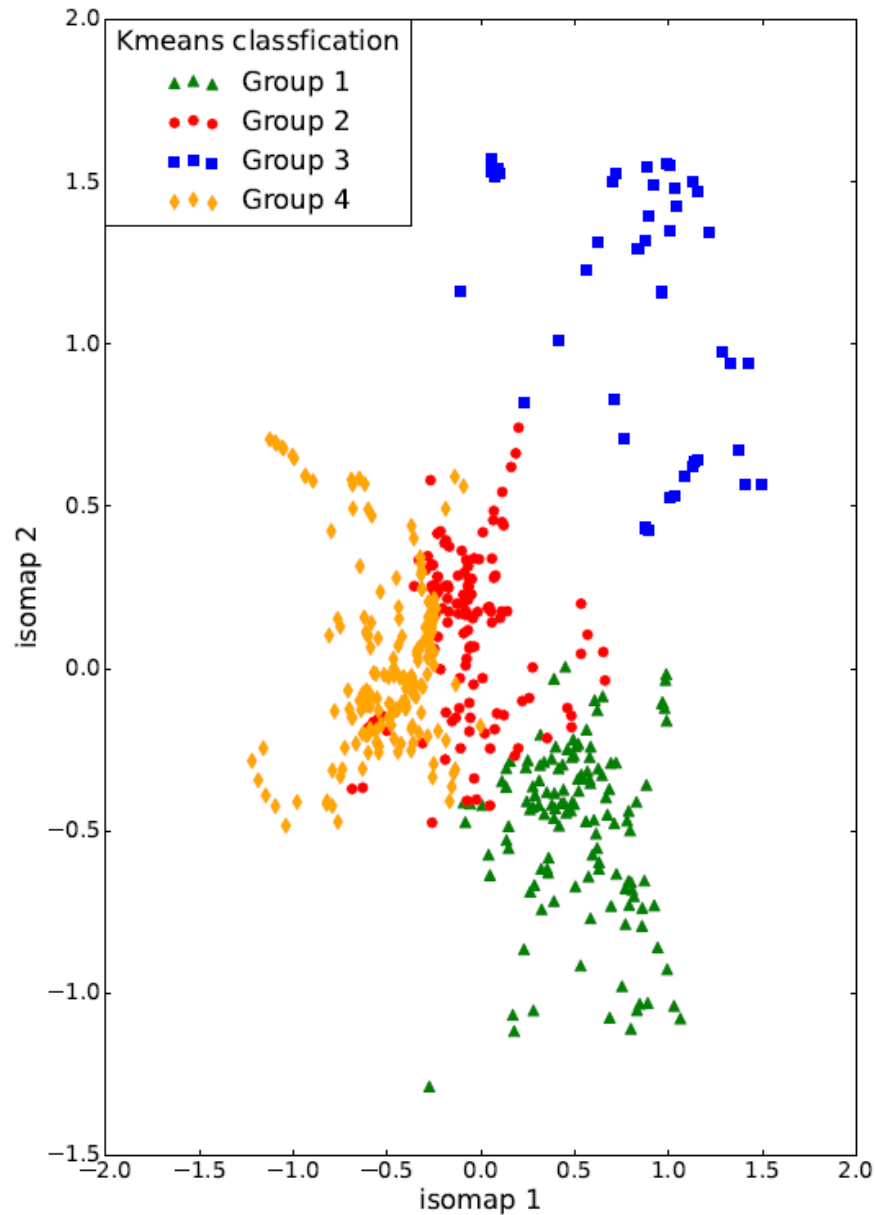
Number of clusters as input

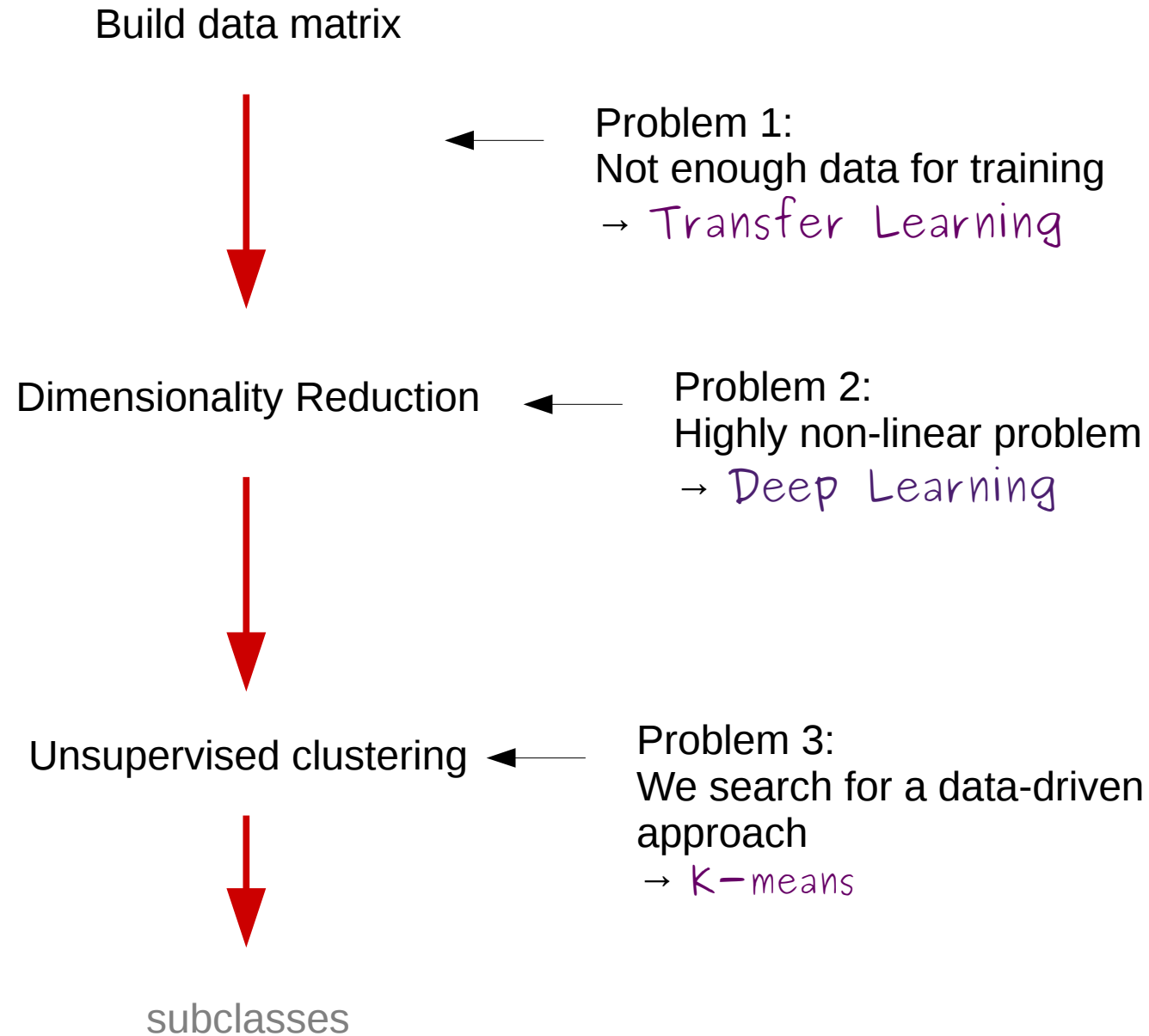




2D visualization of 4D Deep Learning parameter space

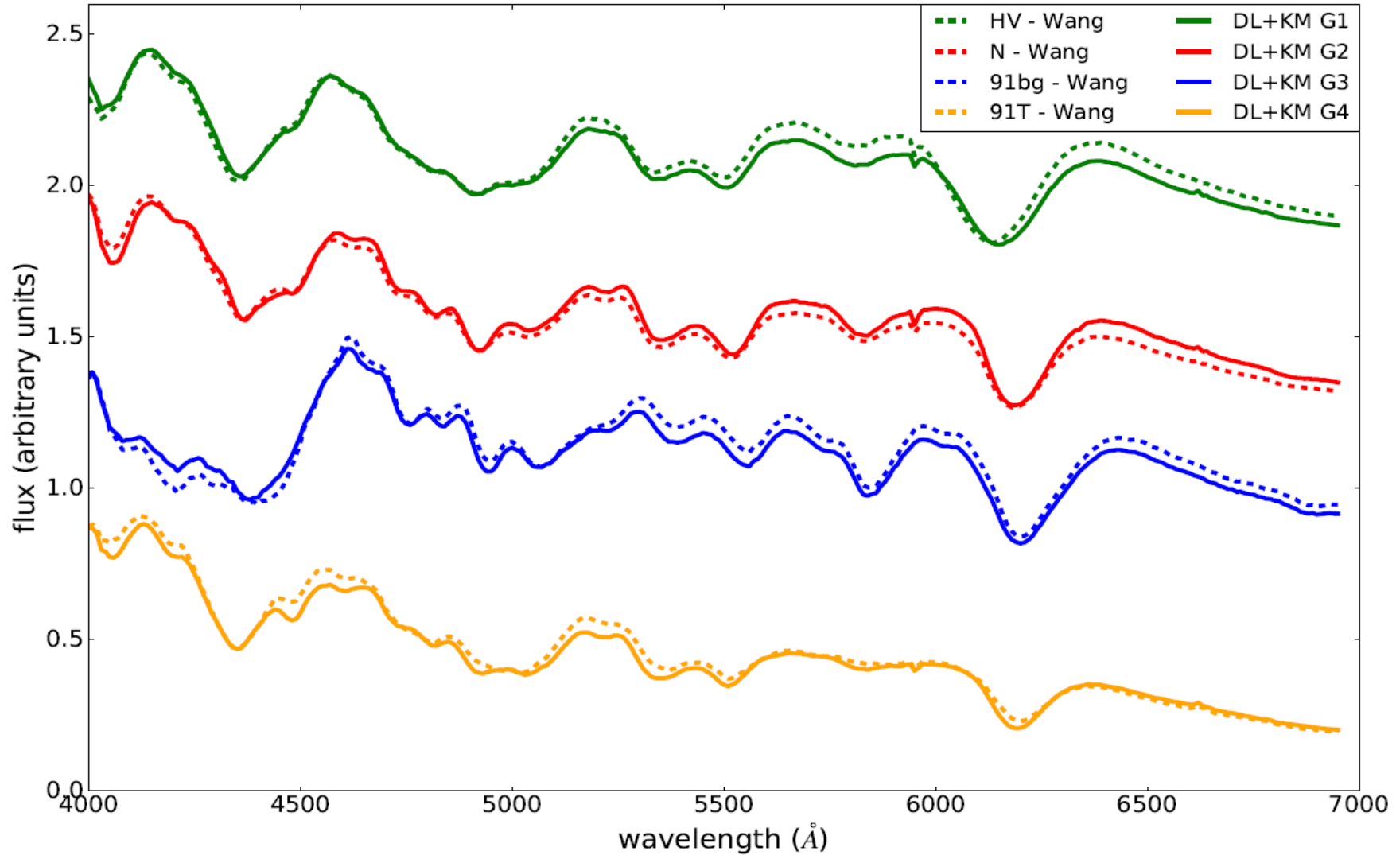
Results from K-means

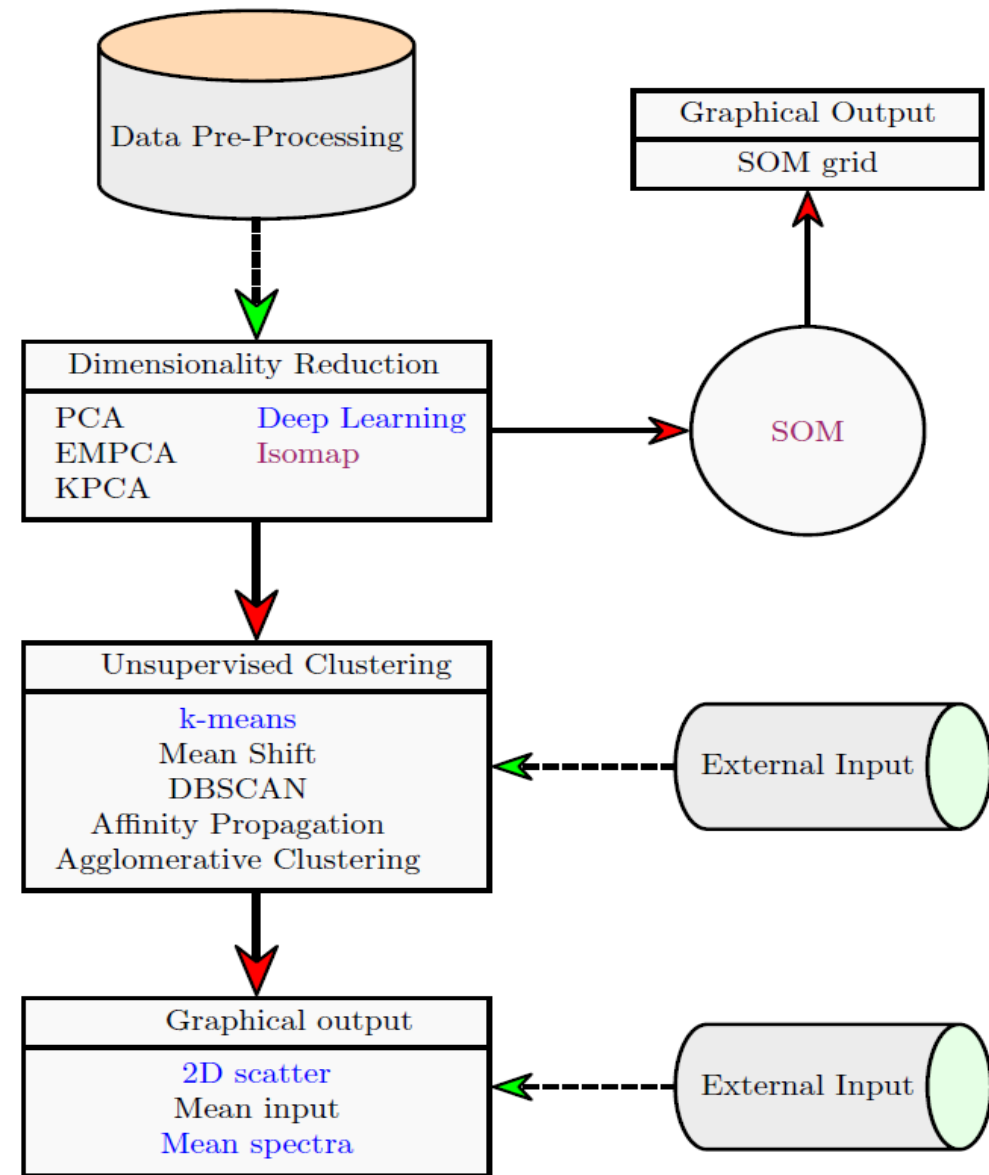




Mean spectra by ML x human

K-means with 4 groups x Wang





COIN'S
Dracula
TEAM



Michele Sasdelli

Project leader
Liverpool John Moores University, UK
Postdoc - on site
Now in Cortexica



Emille Ishida

U. Clermont-Auvergne, France
Postdoc - on site



Ricardo Vilalta

U. Houston, USA
Senior - remotely



Michel Aguená

Head of software development
U. of Sao Paulo, Brazil
PhD student - remotely



Vinicius Busti

U. of Sao Paulo, Brazil
Postdoc - remotely



Hugo Camacho

U. of Sao Paulo, Brazil
PhD student - remotely



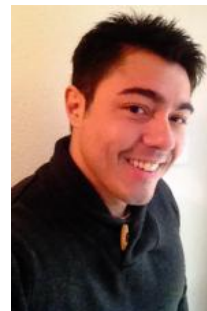
Arlindo Trindade

U. of Porto, Portugal
PhD student – on site
Now in Rolls-Royce



Fabian Gieseke

Radboud University Nijmegen, Netherlands
Postdoc – remotely
Now in Denmark



Rafael de Souza

ELTE, Hungary
Postdoc – on site



Yabebal Fantaye

U. Rome Tor Vergata, Italy
Postdoc – on site
Now in South Africa



Paollo Mazzali

Liverpool John Moores University, UK
Senior - remotely
MAESTRO, Marseilles – June/2017

What now?

Is there a future in this kind unsupervised learning?

What now?

Is there a future in this kind unsupervised learning?

Maybe - if the data is abundance and of good quality

Human classification is not based on current measurement alone

What now?

Is there a future in this kind unsupervised learning?

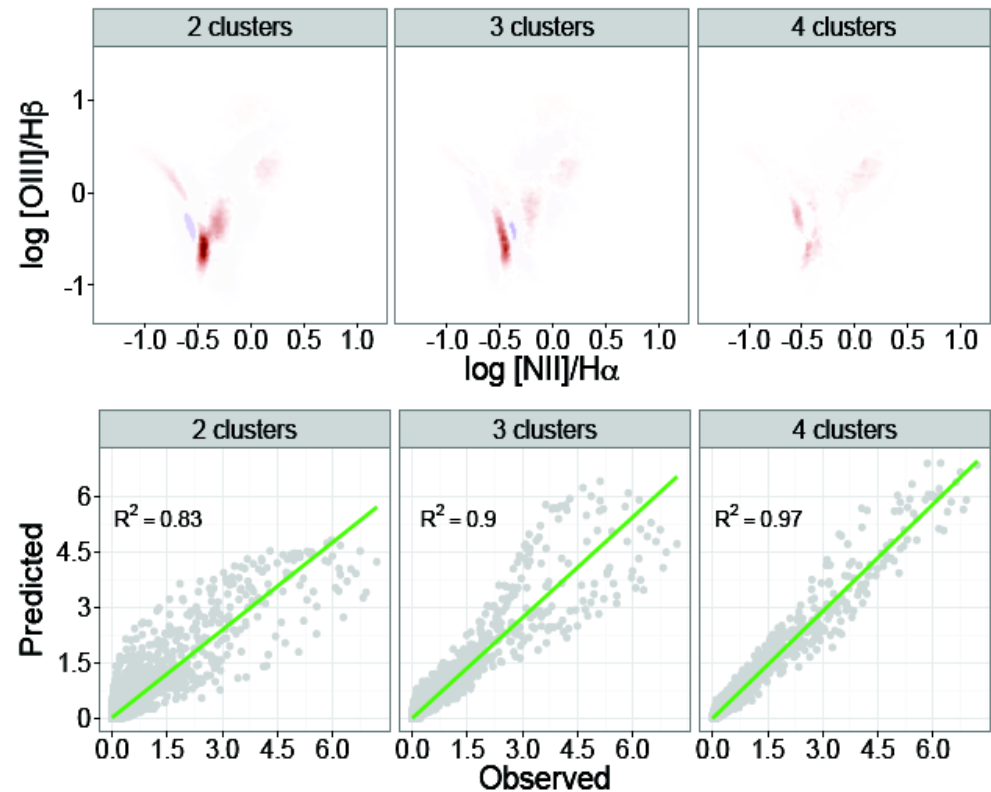
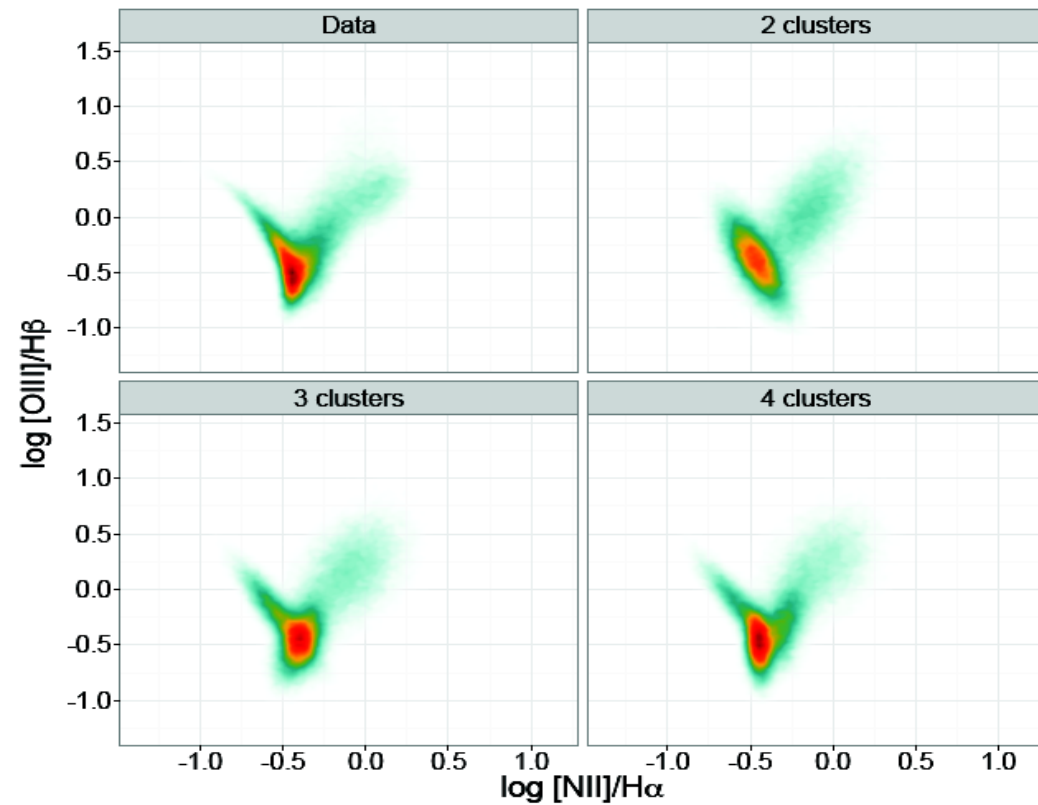
Maybe - if the data is abundance and of good quality

Human classification is not based on current measurement alone

Number of clusters - statistics might help

External clustering validation

CRP #4: External Cluster Validation



What now?

Is there a future in this kind unsupervised learning?

Maybe - if the data is abundance and of good quality
Human classification is not based on current measurement alone

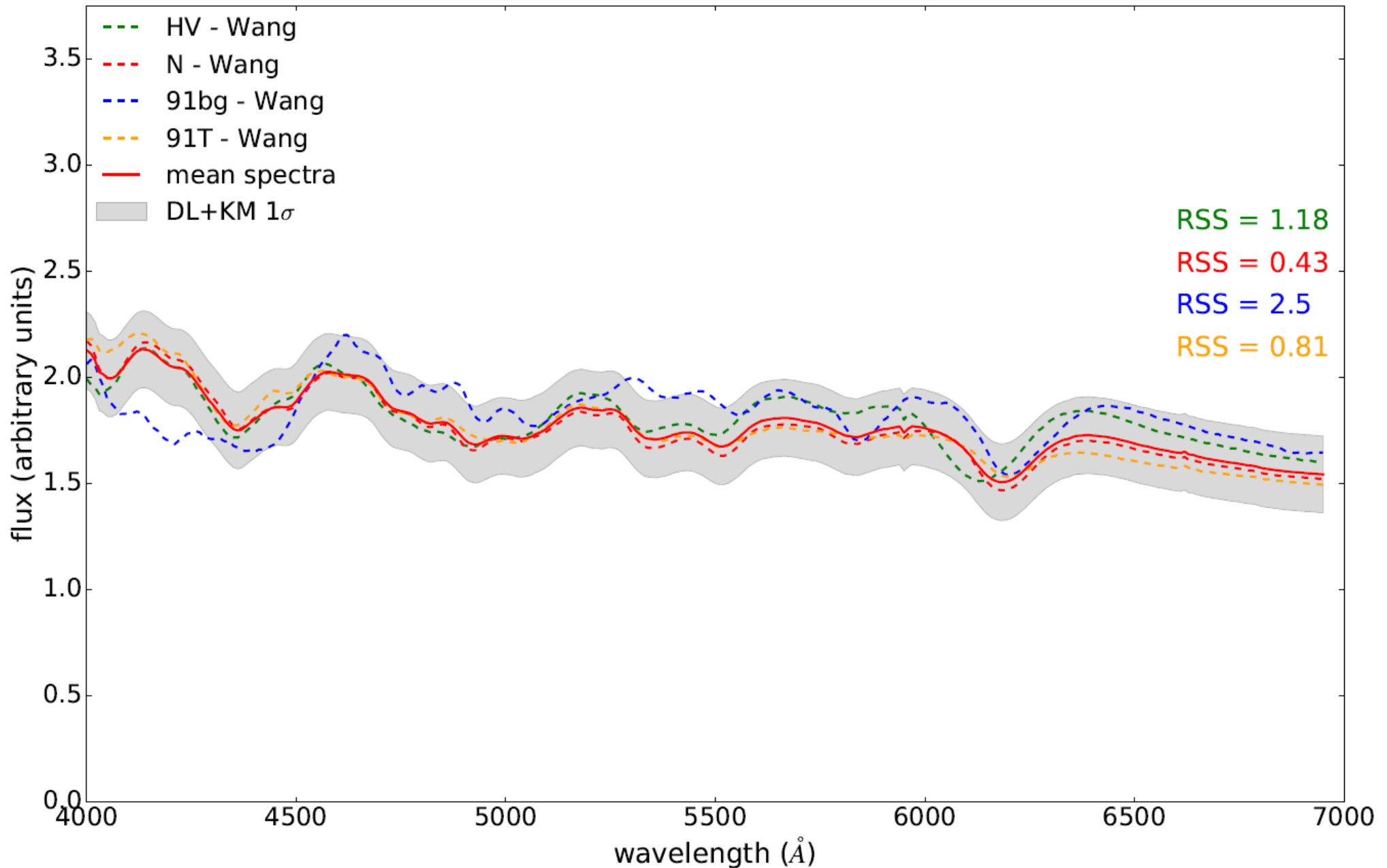
Number of clusters - statistics might help
External clustering validation

Can data guide the theoretical modeling?
To be continued...



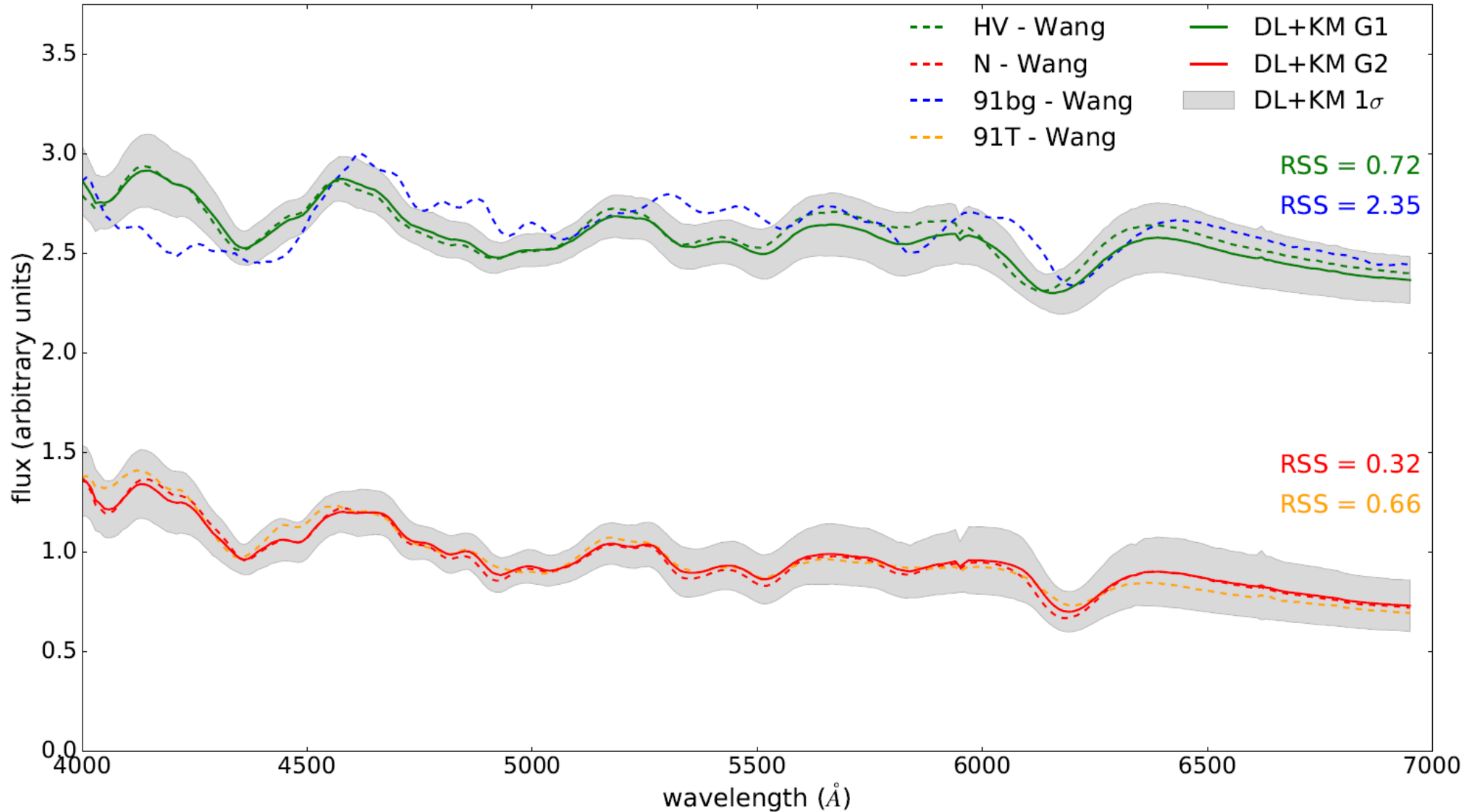
Mean spectra by ML x human

K-means with 1 group



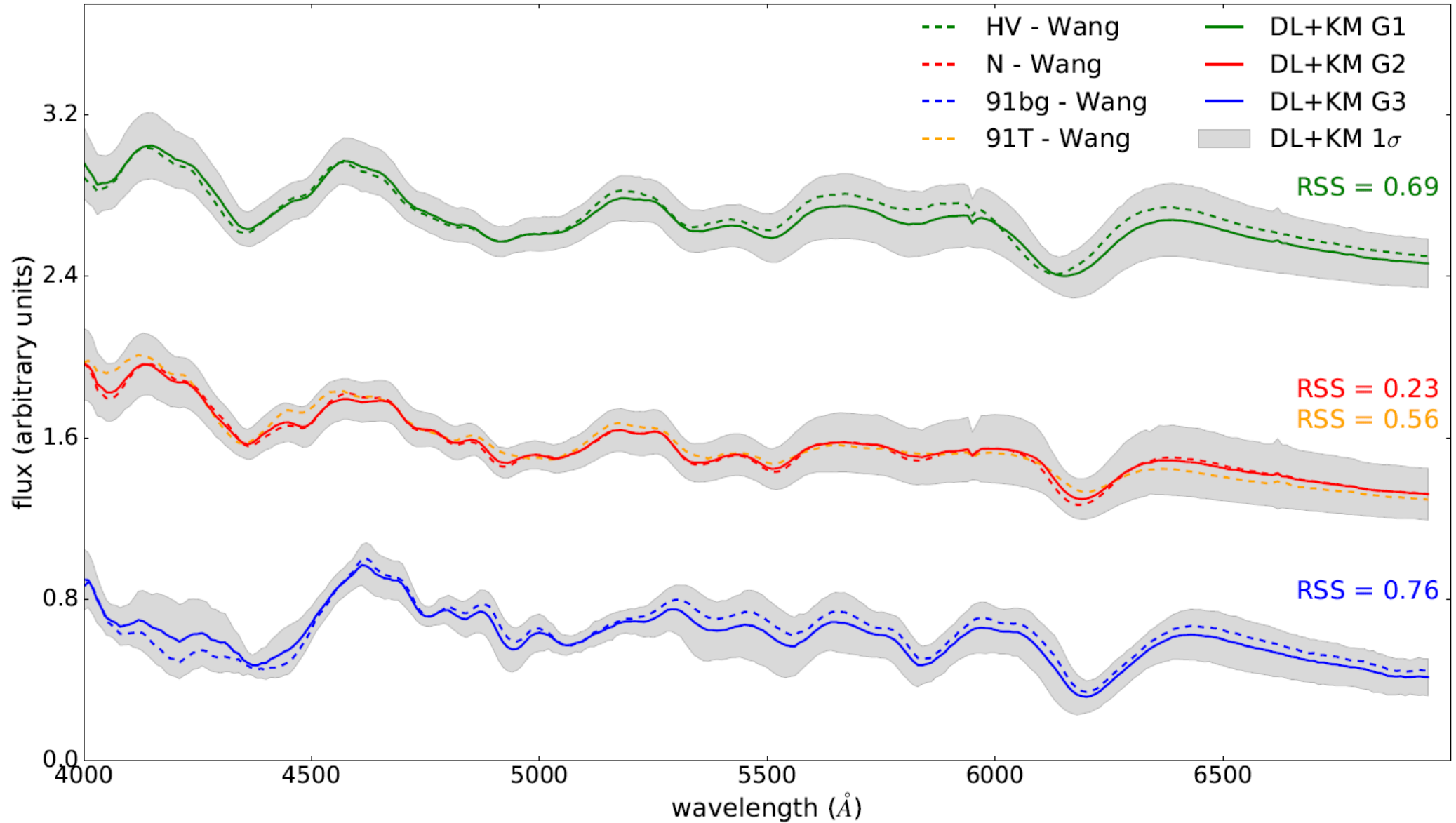
Mean spectra by ML x human

K-means with 2 groups



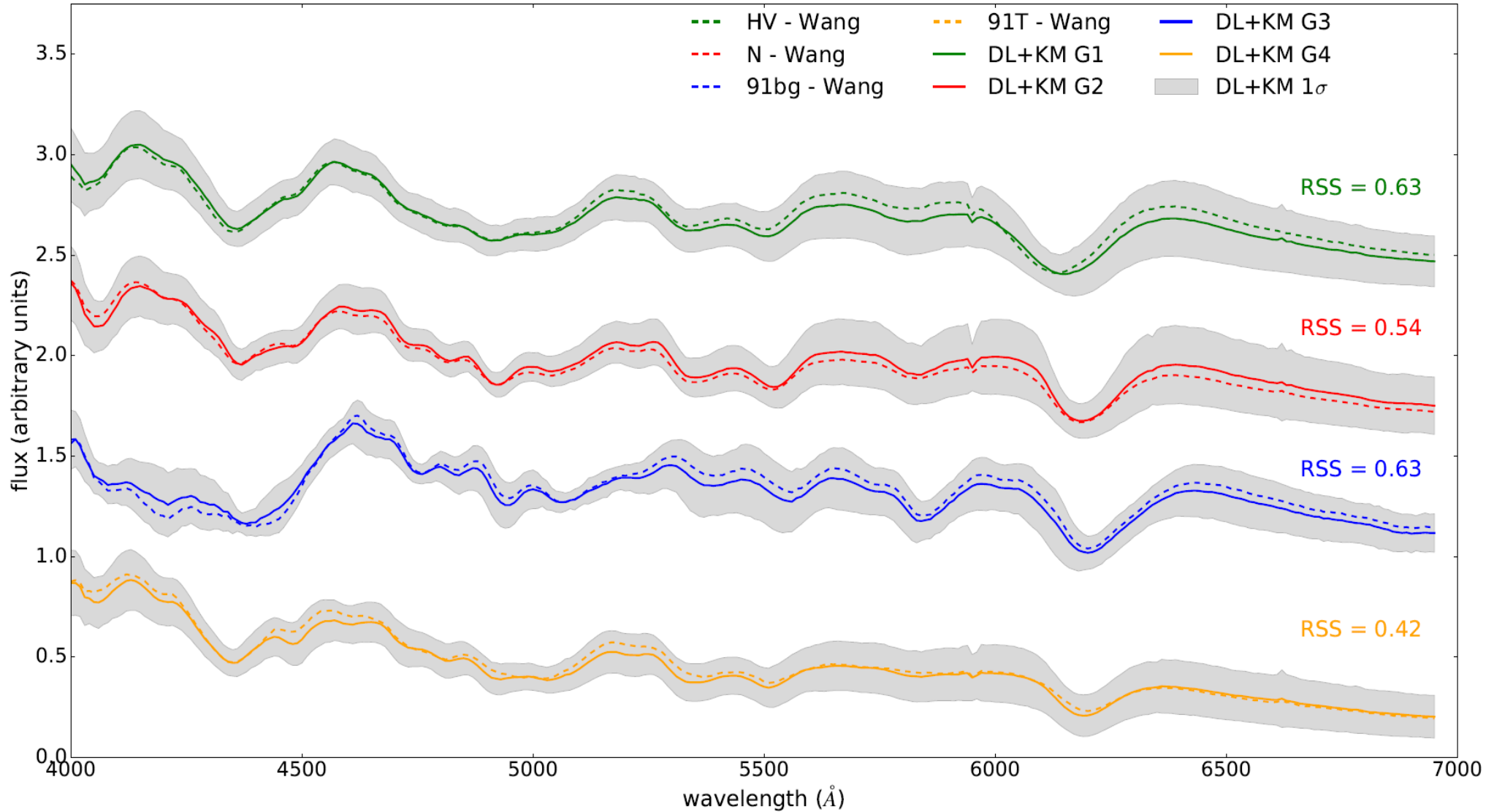
Mean spectra by ML x human

K-means with 3 groups



Mean spectra by ML x human

K-means with 4 groups



In CRP #3
Budapest, 2016

COIN Residence Program #4

Clermont Ferrand, 20-27 August 2017



IAA facebook page

International Astrostatistics Association

Background image by ESO

Organização sem fins lucrativos em Milão
5.0 ★★★★★

Dicas da Página Ver tudo

Como criar publicações eficientes
Publicações pequenas, visuais, criadas para o público certo têm mais êxito.

COIN on twitter

IAA-COIN
@iaa_coin FOLLOWS YOU

COIN promotes the development of novel statistical tools for astronomy. #rstats #astrostatistics #cosmology #python #datascience #astronomy #bigdata

Worldwide
goo.gl/alvZYA
Joined June 2015

Tweet to Message

Registrations open in July

Press F11 to exit full screen

STAT
ASTRO

School of Statistics for Astrophysics 2017: Bayesian Methodology
9-13 October 2017, Autrans (France)

Login

Deep Learning + Self-Organized Maps

Investigating the parameter space

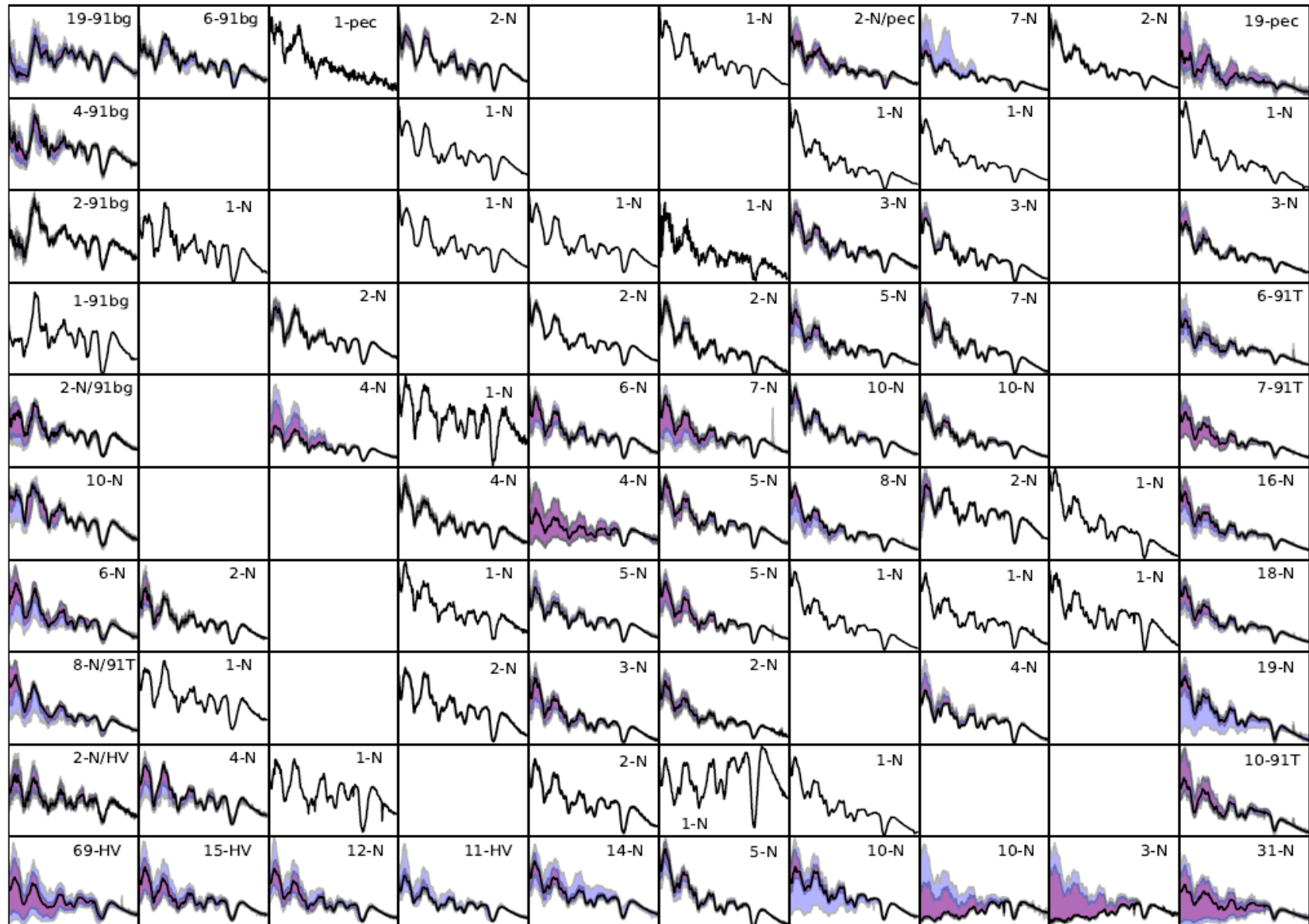


Figure 7. A self organized map of SN Ia spectra at maximum, constructed from the Deep Learning 4-dimensional feature space. Black lines correspond to the mean spectra of each cell, purple and blue bands correspond 1σ and 2σ respectively. Also shown are the number of spectra allocated in each individual cell and the subtype of the majority of SNe populating each cell according to the classification proposed by Wang et al. (2009b). In case there are exactly the same number of objects of different subtypes both labels are shown.

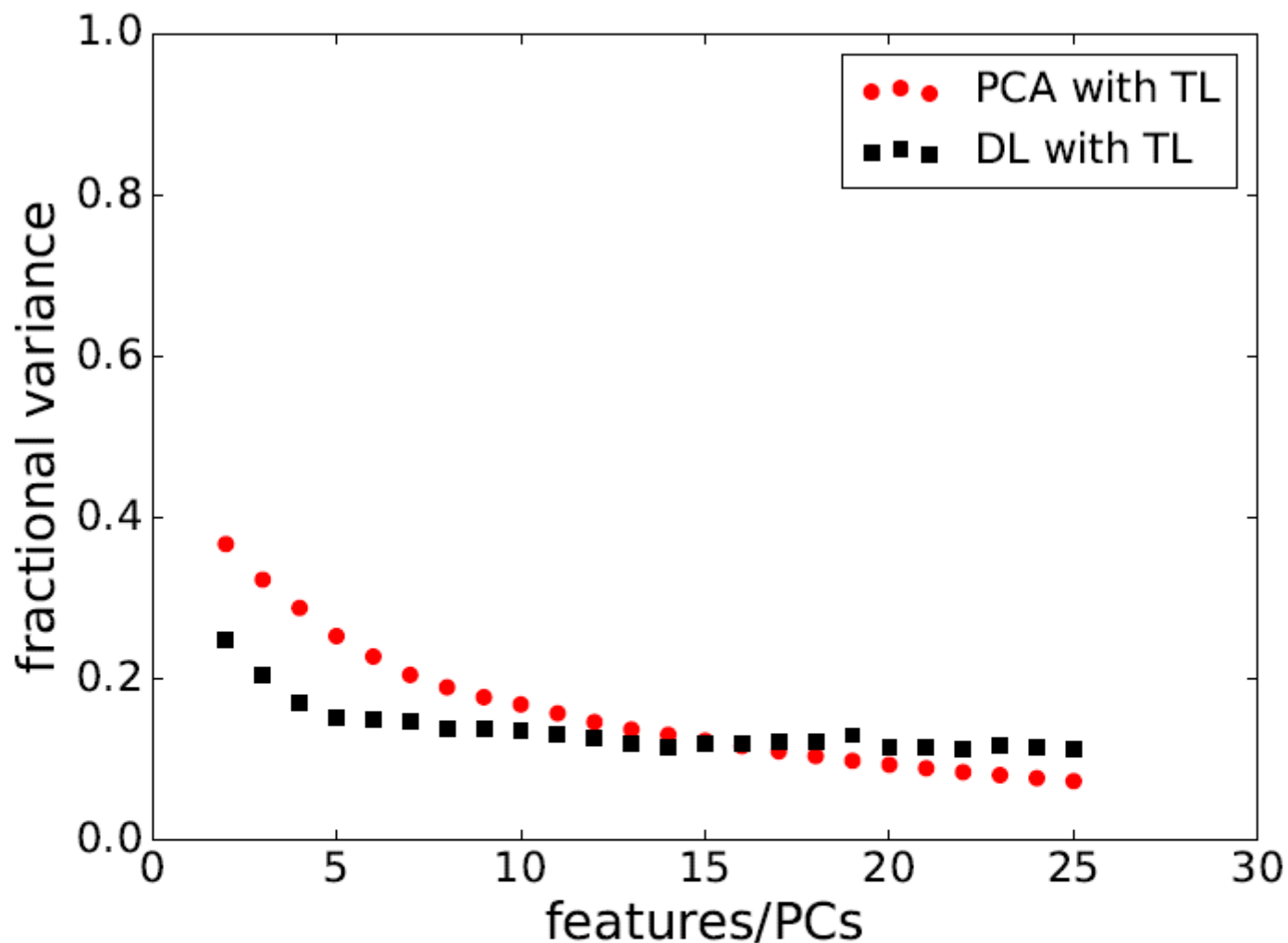
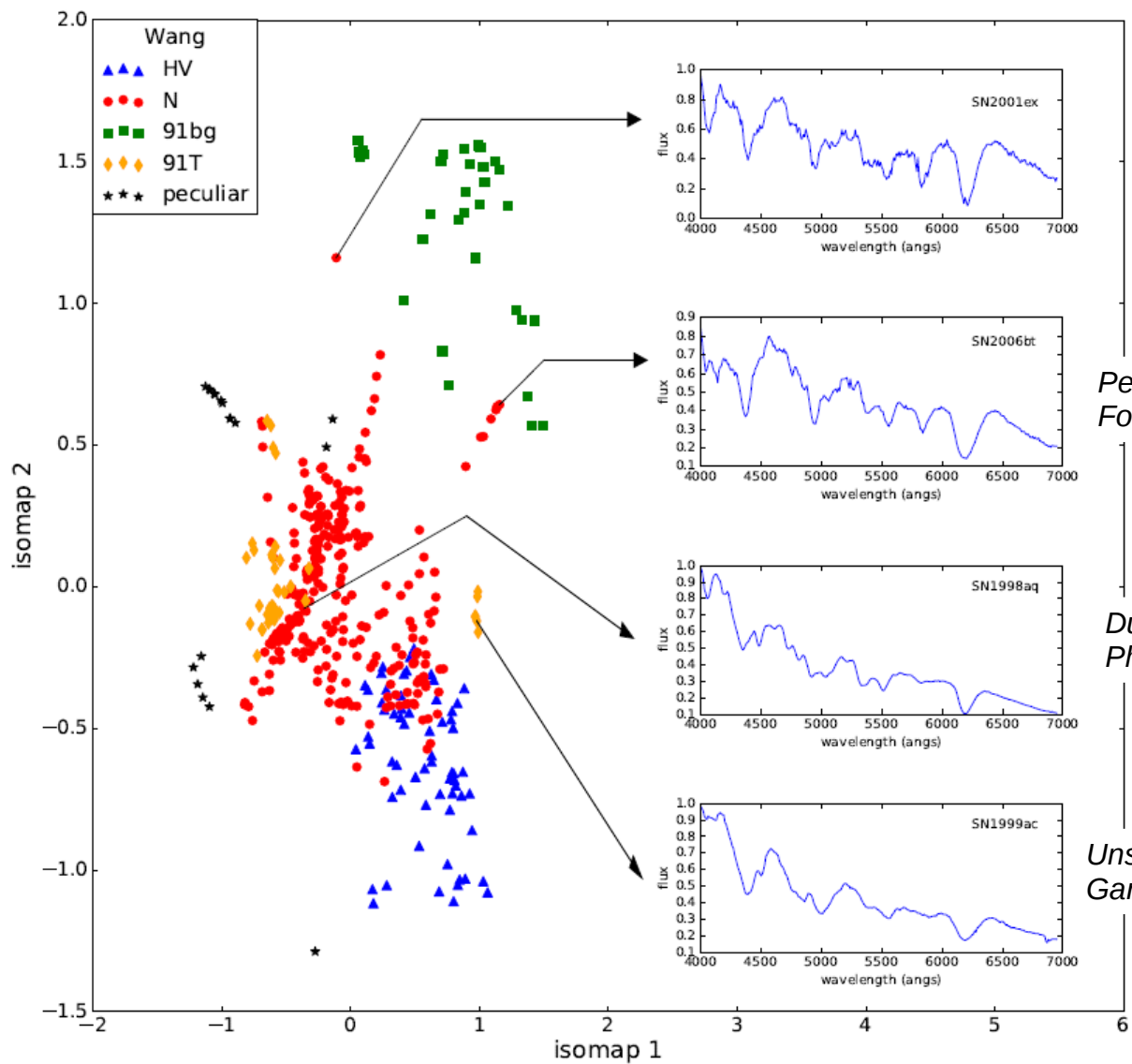
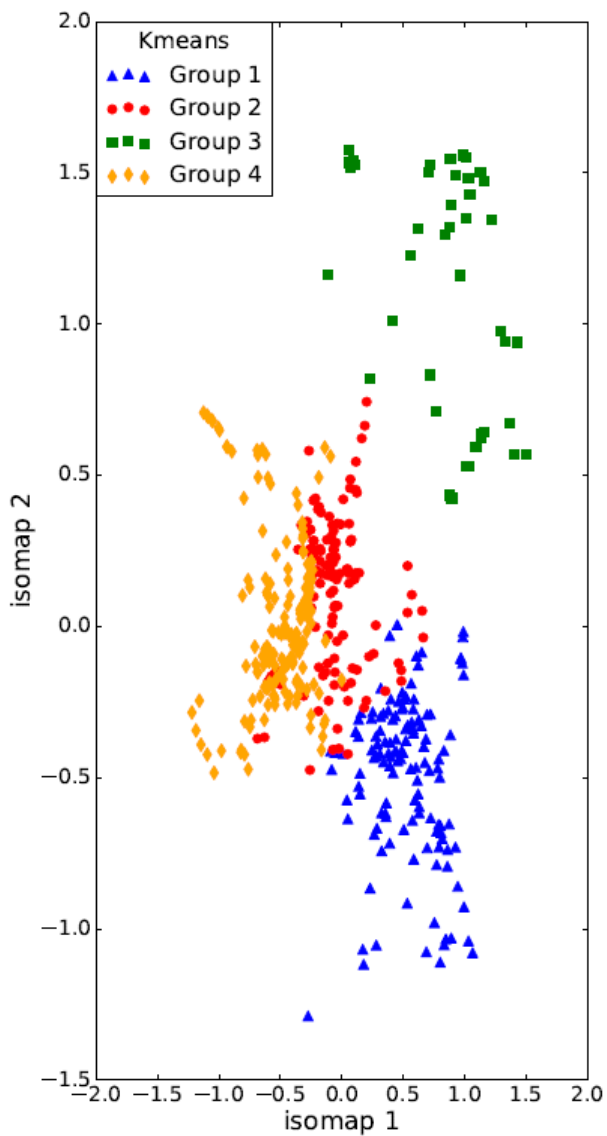


Figure 6. Comparison between Deep Learning and PCA in their capacity to reconstruct the original spectra and robustness to overfitting. Horizontal axis stands for the number of PCs/features and vertical axis shows the deviation between real data and reconstruction.

2D visualization of 4D Deep Learning parameter space

Results from K-means



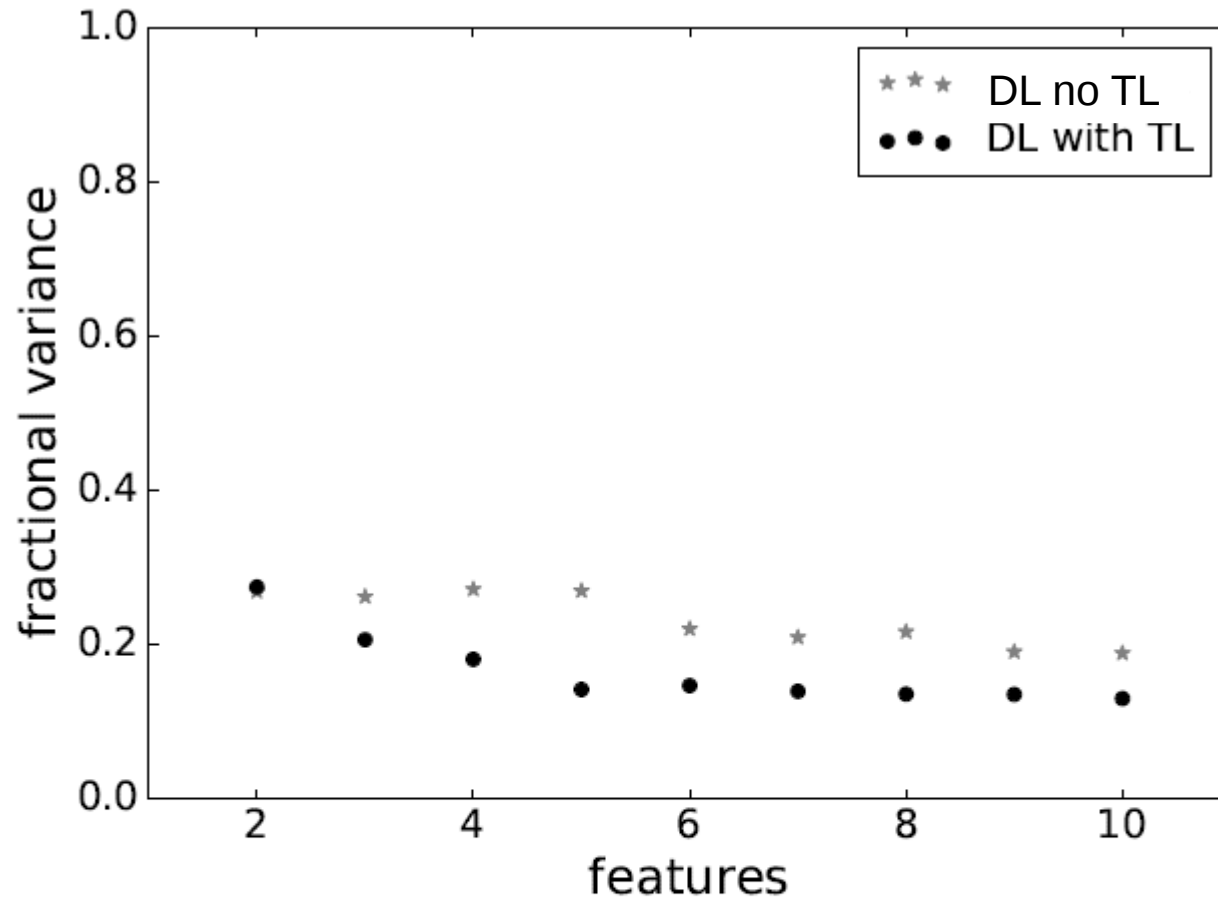


Figure 4. Variance between the deep learning reconstructions and observed SN Ia spectra at B max. Gray crosses correspond to a data configuration using only SNe Ia at maximum (no transfer learning) and black circles denote results from an initial data matrix containing spectra from all epochs (with transfer learning). The horizontal axis stands for the number of features; the vertical axis shows the deviation between real data and reconstruction.